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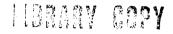
NASA Technical Memorandum 83289

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COMPILATION OF ATMOSPHERIC GAS CONCENTRATION PROFILES FROM 0 TO 50 KM

M. A. H. Smith

MARCH 1982



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LANGLEY RESEARCH IN MER LIFERRY, MASA HAMPTON, VIRGINIA

National Aeronautics and Space Administration

Langley Research Center Hampton, Virginia 23665

INTRODUCTION

The compilation of atmospheric gas concentration profiles described in this report was created at NASA Langley Research Center as a convenient on-line reference data set for line-by-line simulations of atmospheric infrared absorption and emission signals. Profiles from this compilation are also useful as "initial guesses" in iterative procedures for retrieval of gas concentrations from measured atmospheric infrared data. Since a number of researchers in the field of atmospheric remote sensing have informally requested copies of this compilation, the author has decided to release these data in the form of a Technical Memorandum.

In this report, a brief description of the data will be followed by a listing of the data set as it appears on the NASA Langley computer system, and a list of references. The final section of this report contains plots of each of the 52 gas concentration profiles included in the compilation. The author acknowledges the assistance of Ms. Pamela L. Rarig, of the Systems and Applied Sciences Corporation, who developed the plotting program used to produce the figures in this report.

N82-22822 #

DESCRIPTION OF THE DATA

The first 28 gases on the compilation are those for which spectroscopic data are given on the AFGL (Rothman, 1981; Rothman et al., 1981) and GEISA (Chedin et al., 1980, 1981) atmospheric absorption line parameter compilations. Concentration profiles for these gases are listed in order according to their AFGL integer gas codes. Since this gas profile compilation has been developed primarily as a data base for simulation of atmospheric absorption or emission, profiles are included for all the major infrared-active species (H_2O , CO_2 , O_3 , N_2O , CO, CH_4 , O_2), even though some of these profiles may not be of great interest for photochemical studies. The fundamental collision-induced bands of both N_2 and O_2 appear quite strong in tropospheric and lower-stratospheric spectra (Rinsland et al., 1981, 1982), and the atmospheric N_2 profile has also been included in this compilation. PH3 is not known to exist in the unpolluted terrestrial atmosphere, but its line parameters are included in the GEISA tape and are used in analysis of atmospheric spectra of the outer planets. In this compilation, the terrestrial PH3 volume mixing ratio has been set to a negligible value (1.0×10^{-20}) .

The last 24 gases in this compilation are species which have either been measured in the stratosphere or are considered to be important sources, sinks or intermediate reservoirs for photochemically active species in the stratosphere (Hudson and Reed, 1979; Hudson et al., 1981). All of these gases have molecular spectra in the infrared, but detailed listings of their absorption line parameters are not yet available. However, the concentration profiles listed here may be used in conjunction with laboratory spectra such as those compiled by Murcray and Goldman

(1981) to estimate the magnitude of atmospheric absorption by these species at various altitudes. These concentration profiles are loosely grouped according to category, ie., hydrocarbons, halocarbons, odd-hydrogen species, odd-nitrogen species, and sulfur-containing molecules.

The data file is written in the form of 80-character card images. For each species there is a header card containing the gas name left-justified in the first 15 columns, and columns 16 through 80 contain a description of the major references for the gas concentration profile. This header card is followed by one to six cards containing the atmospheric volume mixing ratios at 2-km intervals, beginning at 50 km and ending at the ground (0 km).

The gas concentration profiles in this data set have been based on both atmospheric measurements and the results of photochemical models, compiled from an extensive literature search by the author. The majority of references for these profiles are included in the Hudson et al. (1981) report, and a few more recent results are added here. Although for a number of species many measurements and model studies have been reported in the literature, the references on the header cards are limited to the more recent results and to review articles. Unless otherwise stated on the header card, the set of volume mixing ratios for each species represents the "typical" diurnally averaged, seasonally averaged Northern Hemisphere midlatitude gas concentration profile. When both model results and atmospheric measurements were available, the measurements have been weighted more heavily in the determination of the average concentration profile. In some cases where neither measurements nor model results were available over the entire 0-50 km altitude range (e.g. NH3, HI, HCN, C2H4,

CHCl $_2$ F, HNO $_2$, HO $_2$ NO $_2$ and H $_2$ S), the author has extrapolated the concentration profiles based on available information such as spectroscopic upper limits, expected ground-level concentrations or estimated tropospheric lifetimes.

The average gas concentration profiles presented here, while useful for "ball-park" estimates of atmospheric absorption or emission or for initialization of retrieval algorithms, may be quite different from concentration profiles actually measured at specific times and locations on the Earth, particularly in the case of diurnally varying species. The reader interested in simulation of atmospheric absorption or emission for a specific case, for example, a sunset spectrum taken in mid-winter at high latitude, should consult the individual references given here and in the Hudson et al. (1981) report to obtain more detailed gas concentration profiles.

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2.3E-11,	2.5E-11,	3.0E-11,	5.0E-11,	8.0E-11,			
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C 2H6	CRONN AND ROB	INSON(1979),	RUDOLPH ET AL	• (1981)			
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C2H4
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                                  4.7E-12,
                                                8.5E-12,
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     2.8E-11.
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                                                7.4E-11.
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                   1.7E-10,
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     2.8E-10.
                                  2.9E-10.
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                    2.9E-10.
CF3CL
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     1.8E-12.
                   1.98-12.
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                                                2. 1E-12,
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                   3.6E-12,
                                  3.9E-12,
                                                4.2E-12,
                                                              5.0E-12.
   6*5.0E-12
                RASMUSSEN ET AL. (1981), GOLDMAN ET AL. (1981A), LEIFER ET AL. (1981)
CHCLF2
   5*1.0E-15,
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                    3.0E-13.
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CHCLZF
                PENKETT ET AL. (1980) GROUND-LEVEL MEAS. ONLY
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                   1.0E-15,
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CCL4
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   6*1.4E-10
CF4
                FABIAN ET AL. (1981B), PENKETT ET AL. (1981), GOLDMAN ET AL. (1979)
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                   6.1E-11,
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CH3CCL3
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                    4.0E-12,
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C2F4CL2
                FABIAN ET AL. (1981B), SINGH ET AL. (1979)
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C2F5CL
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                    1.3E-12.
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FABIAN ET AL. (19818), PENKETT ET AL. (1981)
C2F6
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                YUNG ET AL. (1980), LOGAN ET AL. (1979), ANDERSON ET AL. (1981)
HD2
      7.8E-10,
                    7.2E-10,
                                  6.6E-10.
                                                6.0E-10,
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                                                              1.8E-12,
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HNDZ
                MURCRAY(1979), PLATT AND PERNER(1980) UPPER LIMITS
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                                                              2.5E-11,
      2.0E-11
HO2ND2
                MURCRAY (1979) UPPER LIMIT AT 26 KM
   26*4.0E-10
NO3
                NOXON ET AL. (1978, 1980), NAUDET ET AL. (1981) NIGHTIME
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      4.0E-10,
                    3.5E-10.
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   6*3.0E-12
                BOUGHNER AND NEALY(1979), HUDSON ET AL.(1981), KING ET AL.(1976)
N205
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                    1.0E-09,
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      7.0E-14
CLOND2
                HUDSON AND REED (1979), MURCRAY ET AL. (1979)
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      1.0E-12.
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                    6.5E-11.
                                  1.8E-10,
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      8.0E-10,
                    8.3E-10.
                                  8.5E-10.
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                                                              4.3E-10,
      2.6E-10,
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                                                              3.2E-11,
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SZE AND KO (1980), TURCO ET AL. (1981B) MODELS + ALL MEAS.
CSZ
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                                 3.0E-14,
                                                             3.0E-13,
     1.06-12,
                   3.0E-12,
                                 1.0E-11,
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     4.0E-11,
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                                 5.0E-11,
                                               5.6E-11,
                                                             6.2E-11,
     7.0E-11
HZS
                RODHE AND ISAKSEN(1980), JAESCHKE ET AL. (1980), SZE AND KO(1980)
  10*1.UE-15,
                                 1.08-15,
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     1.0E-15,
                   1.0E-15,
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                                               1.0E-12,
     3.0E-14,
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                                                             3.0E-12,
                                                             3.0E-11,
     7.0E-12.
                    1.0E-11.
                                 1.3E-11,
                                                2.0E-11.
     1.0E-10
H2504
                ARNOLD ET AL. (1981), TURCO ET AL. (1981A)
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                                 1.4E-11,
                                               1.9E-11,
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                   2.8E-14,
                                 2.0E-14,
                                                1.5E-14,
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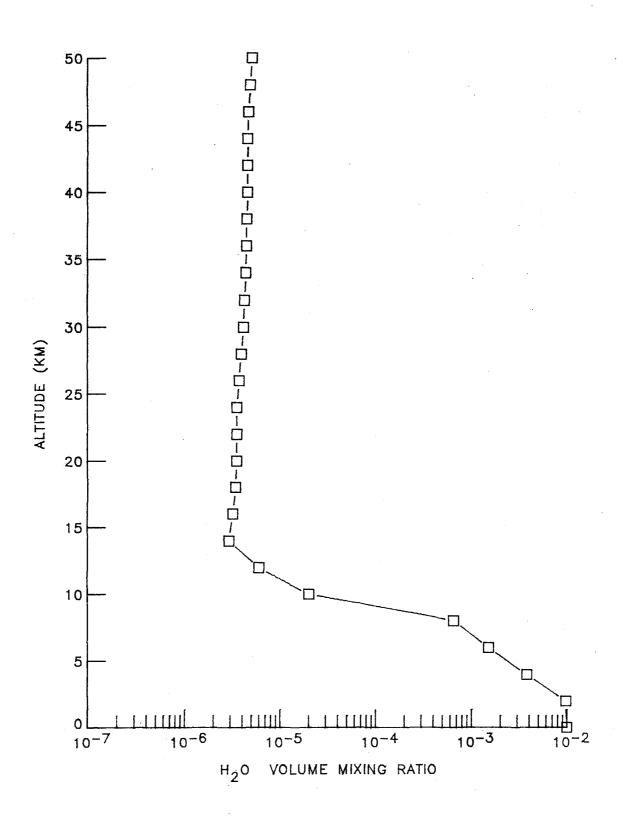
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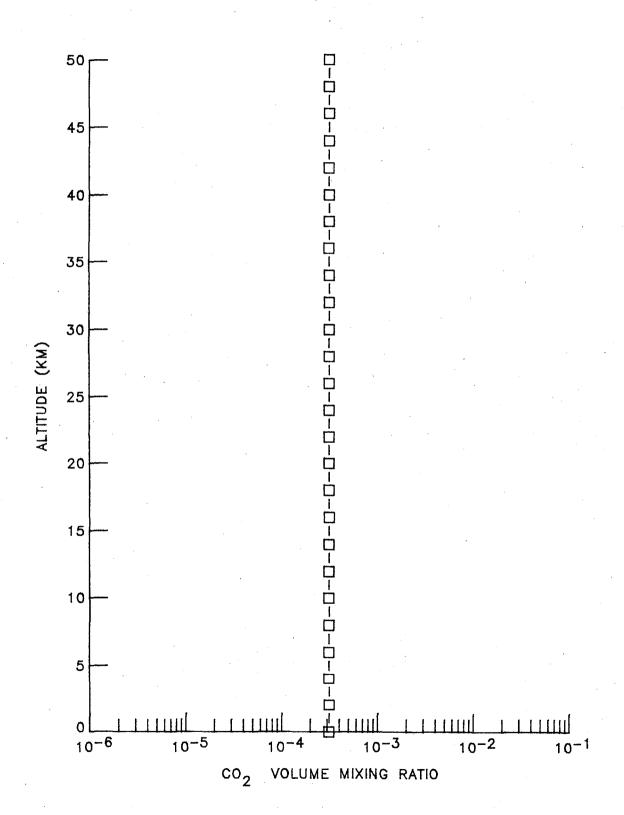
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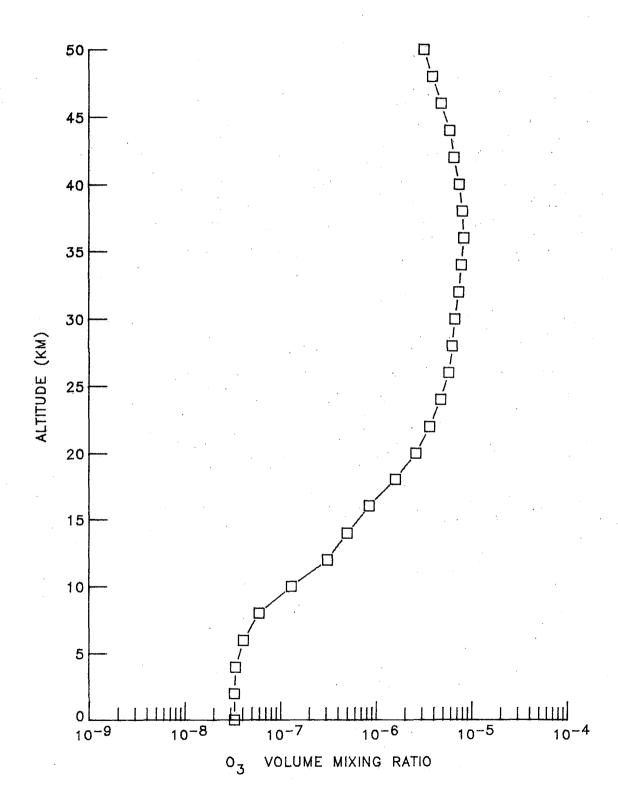
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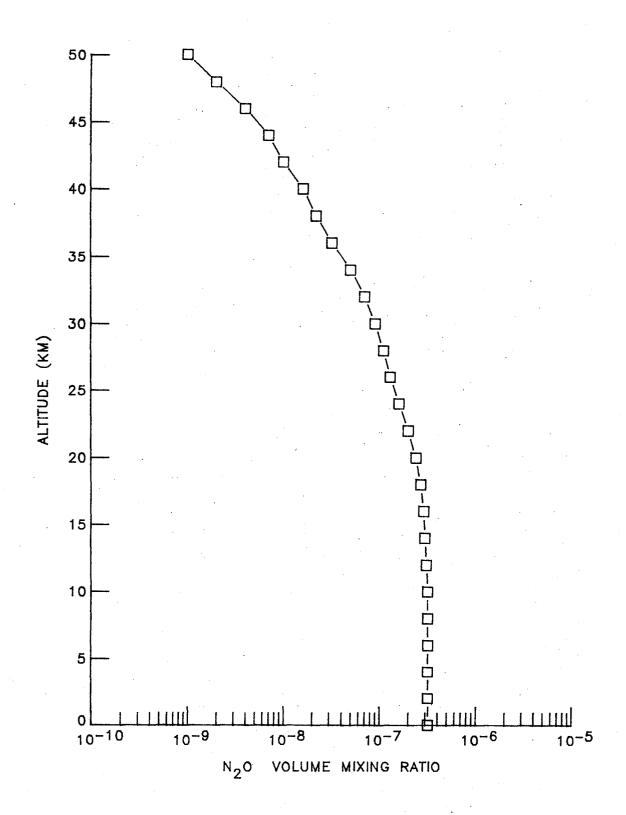
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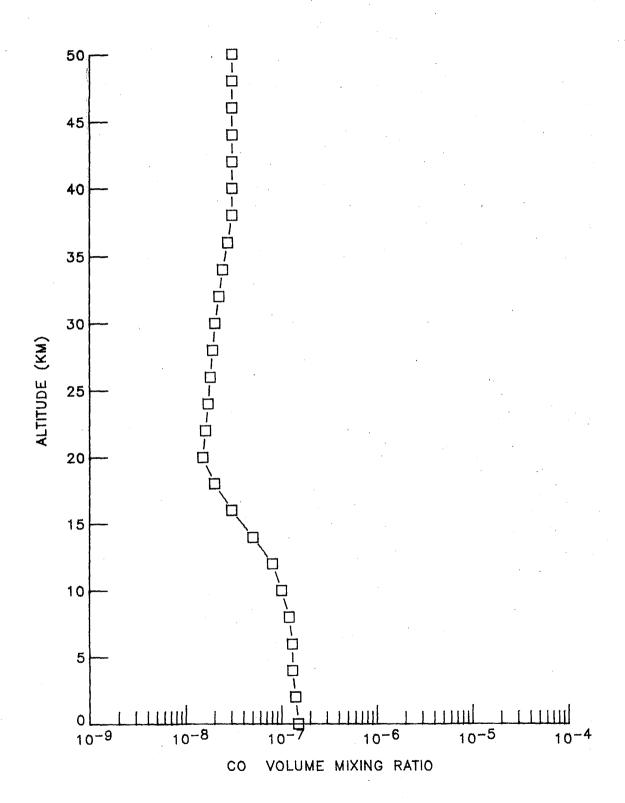
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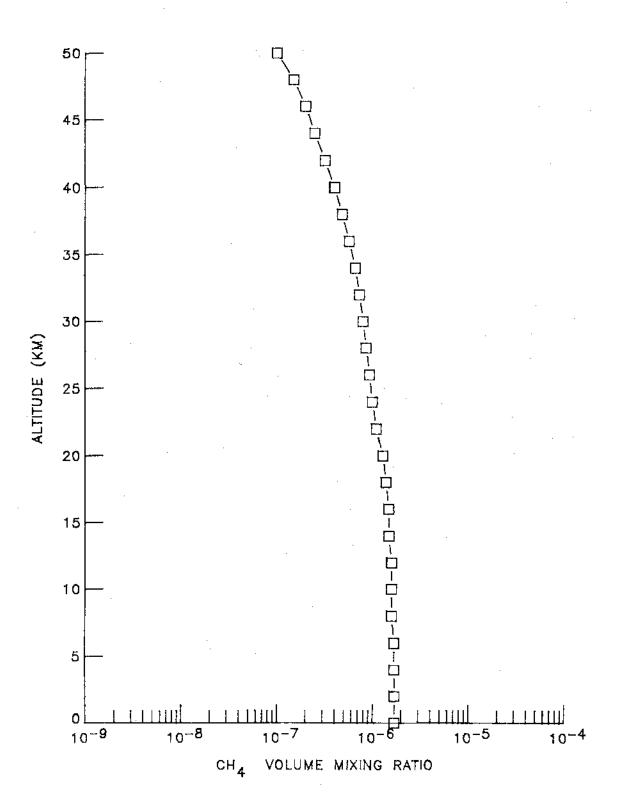


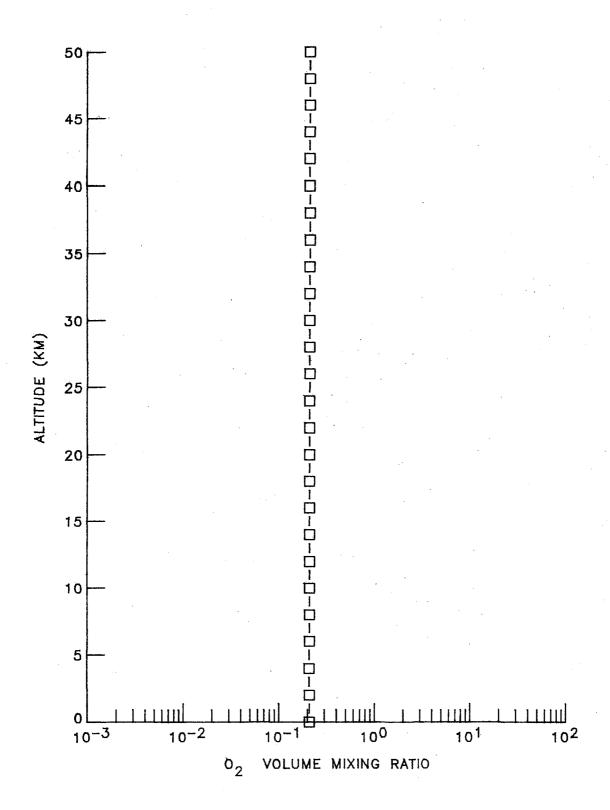


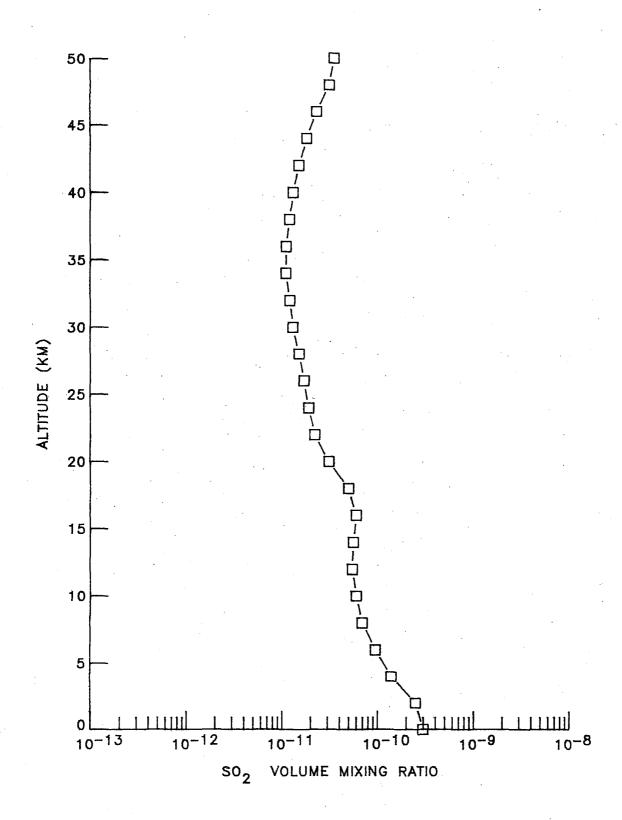


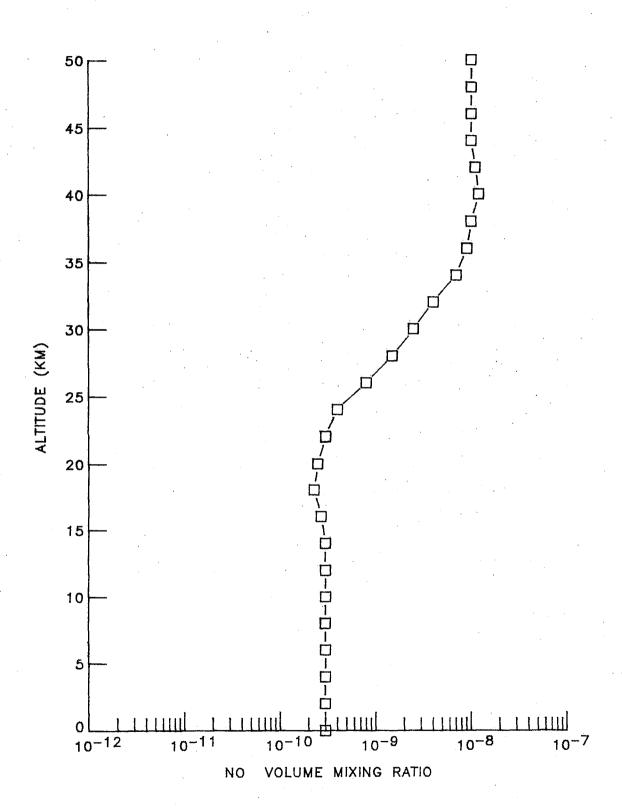


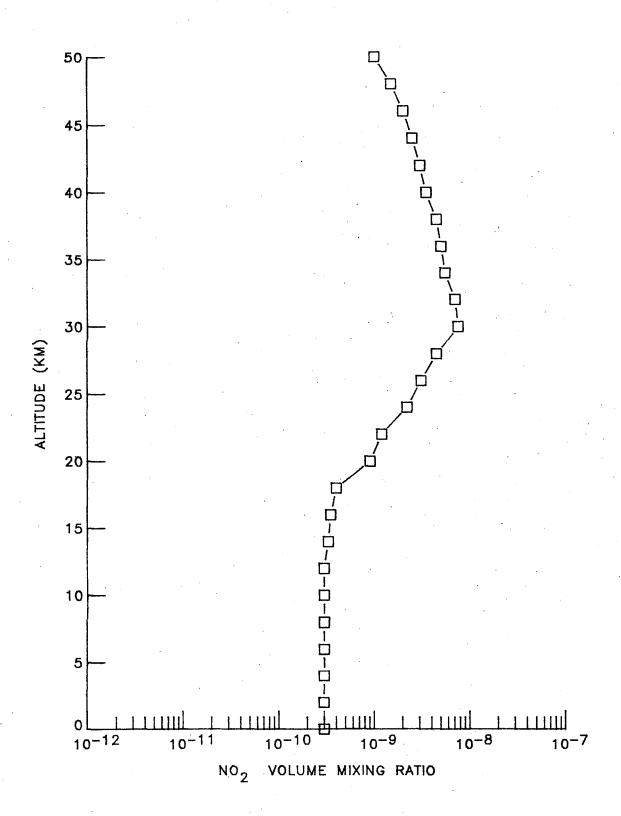


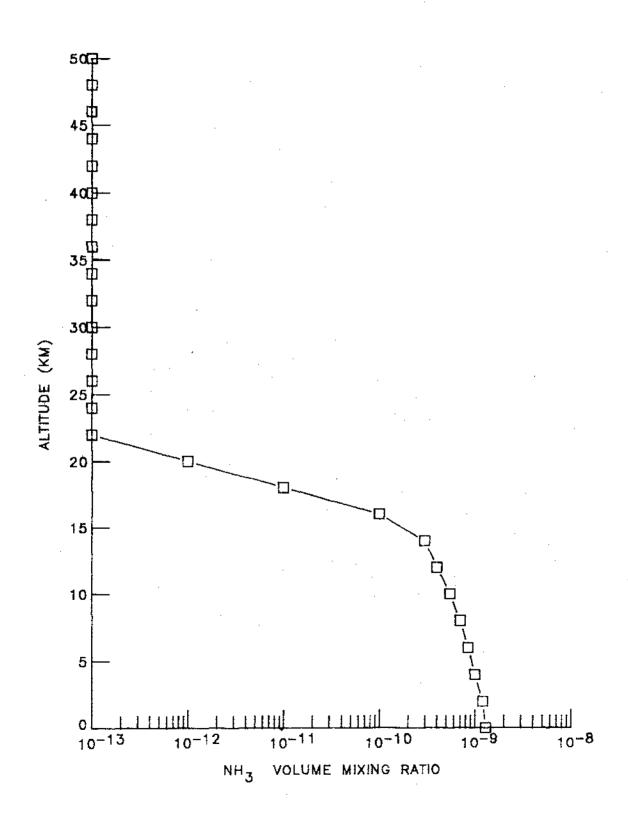


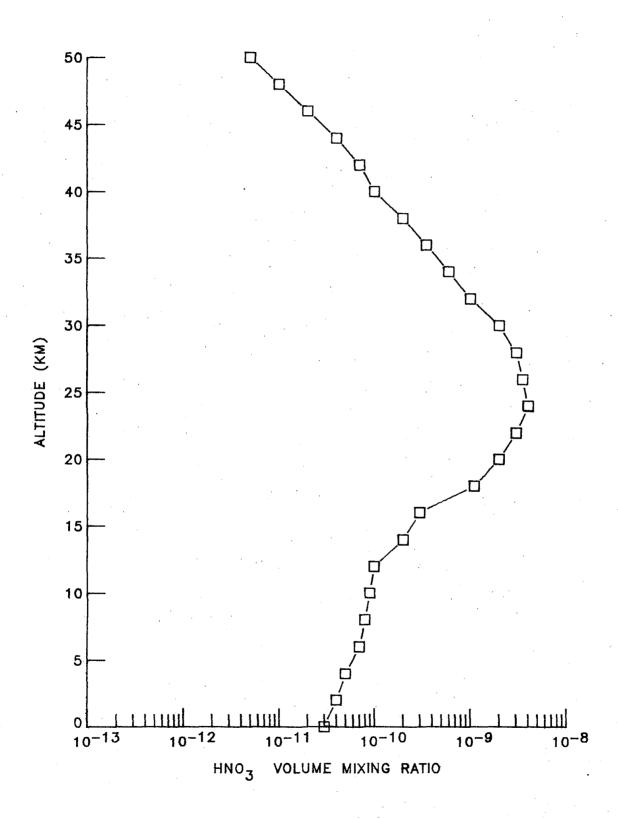


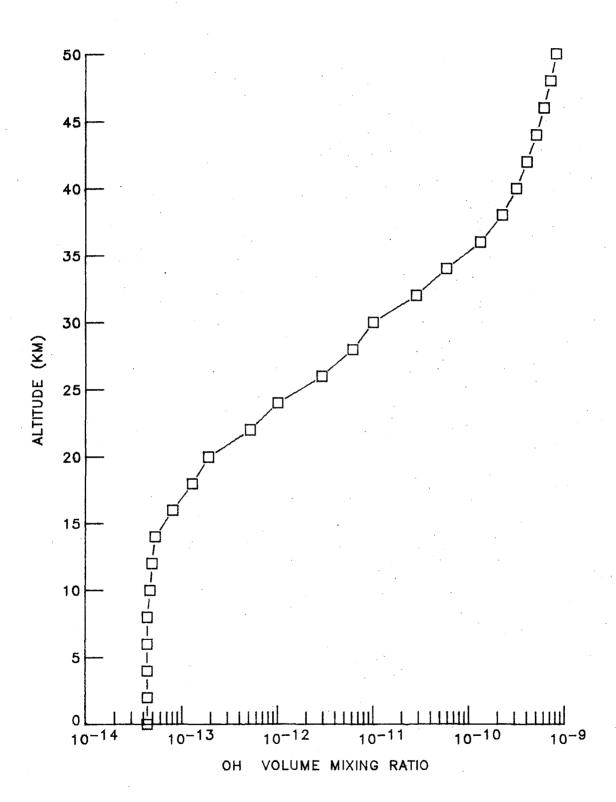


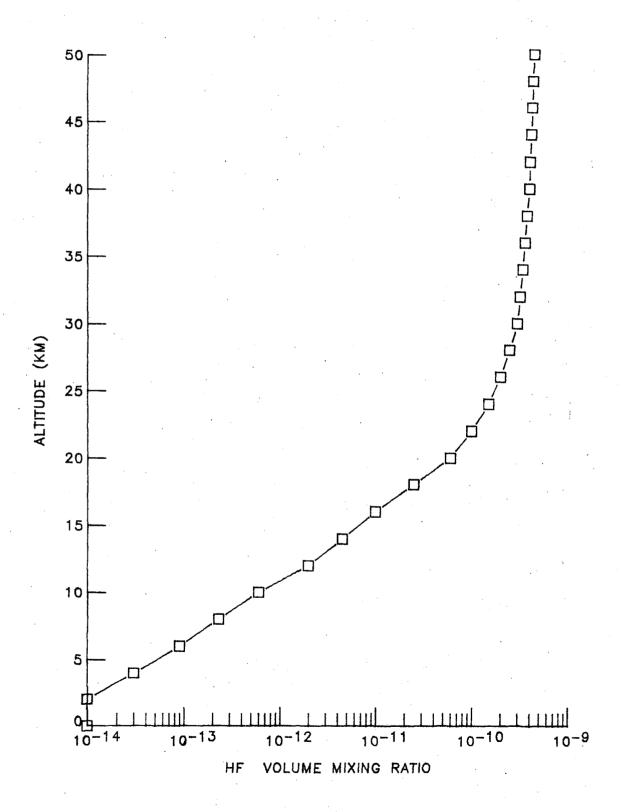


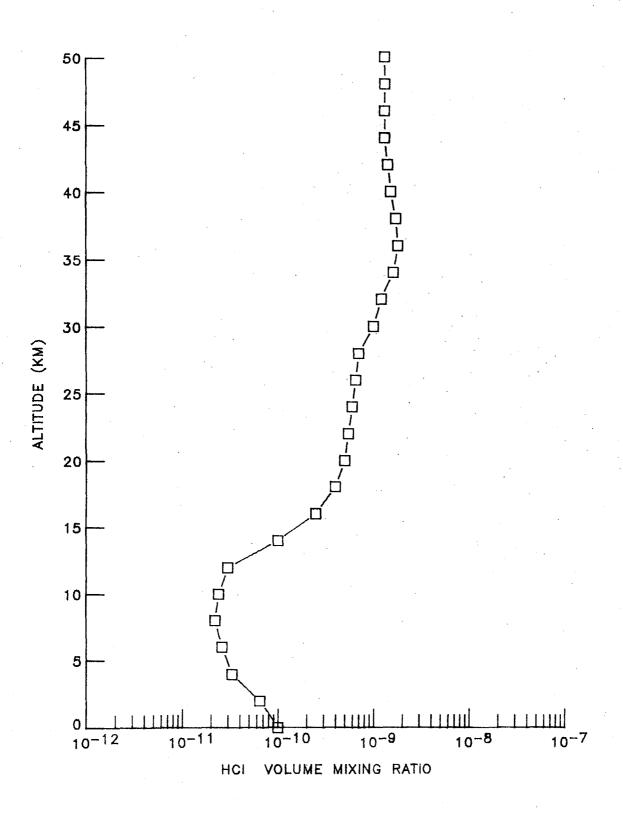


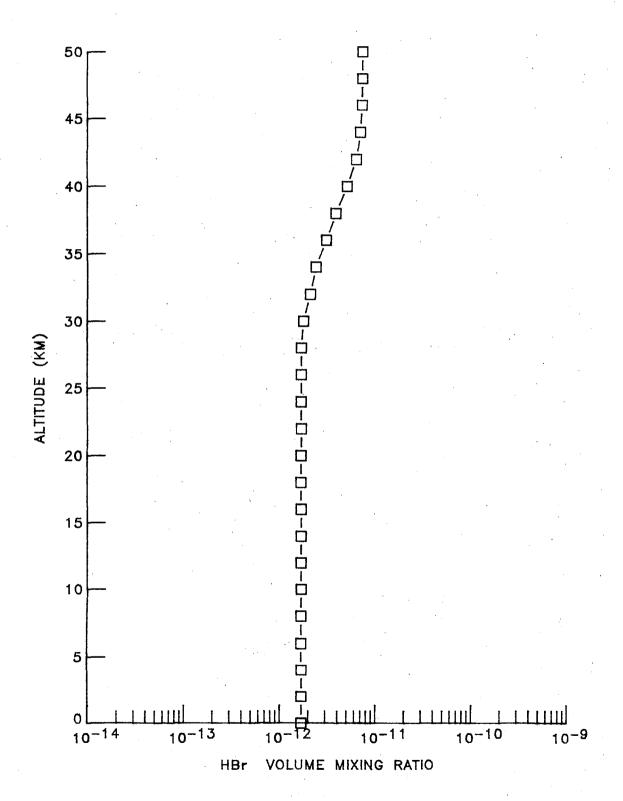


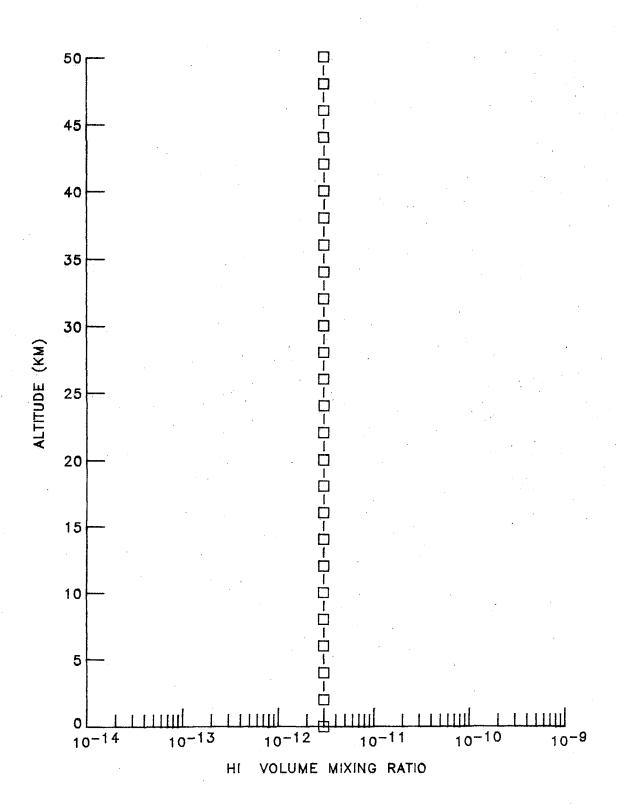


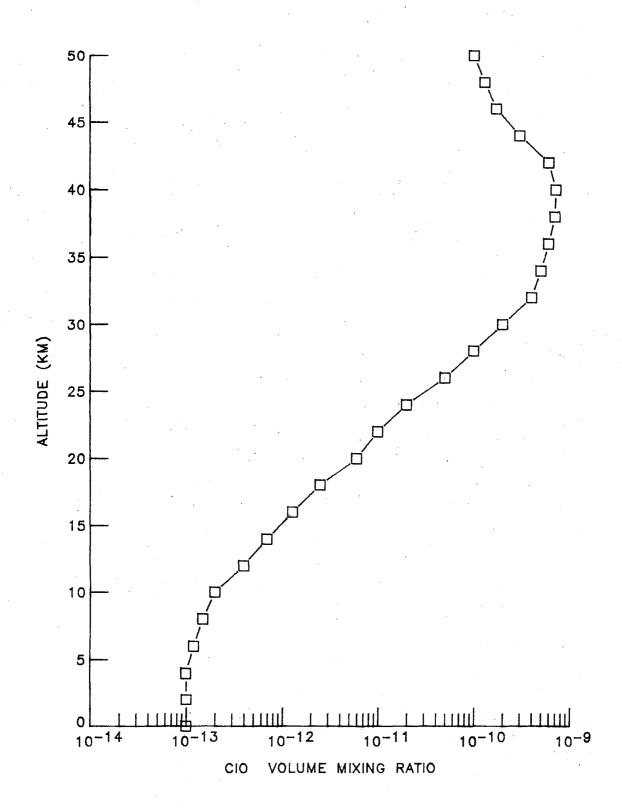


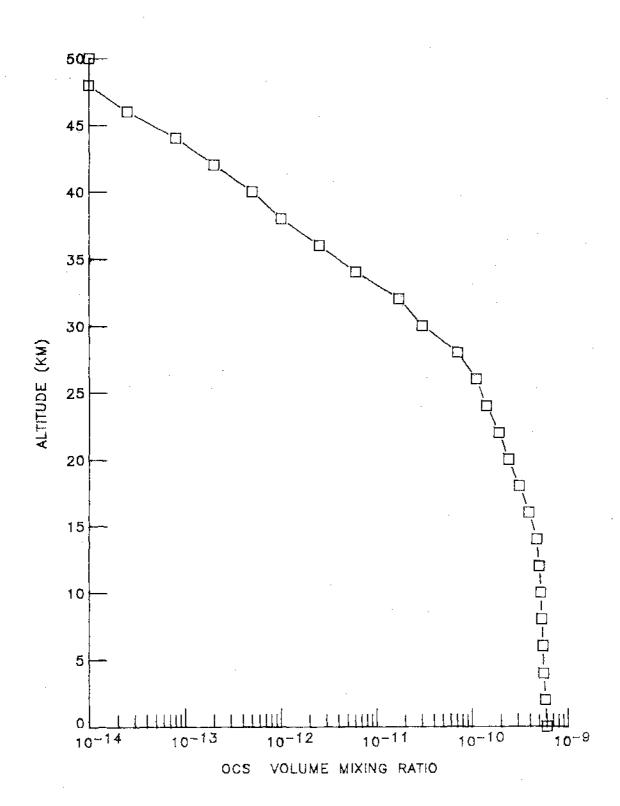


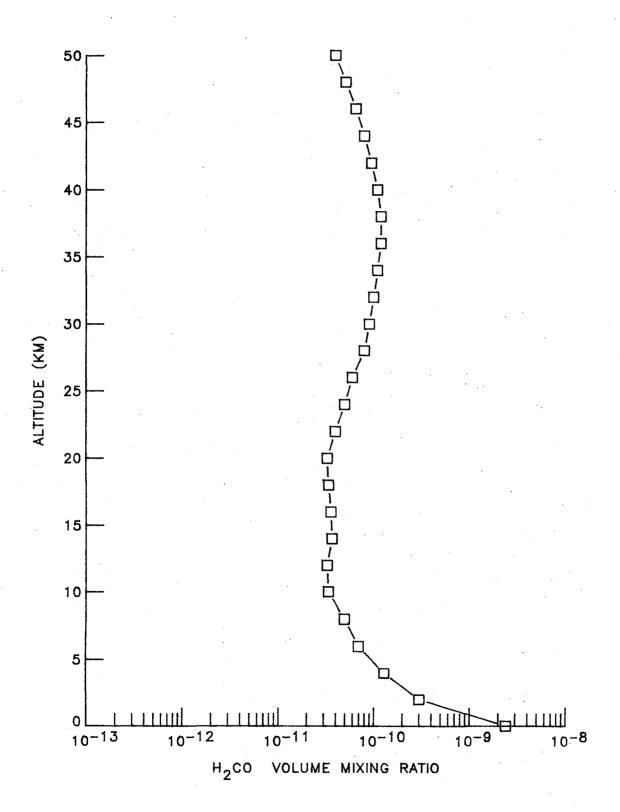


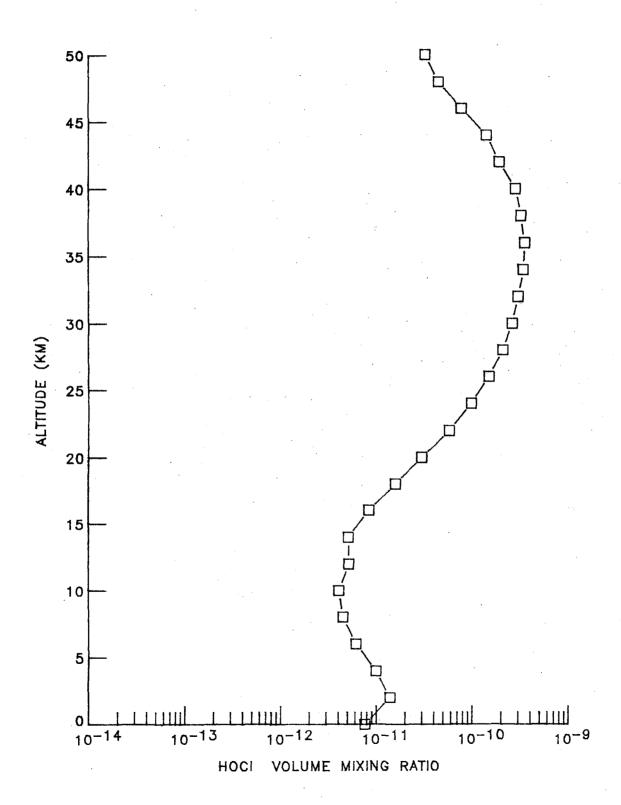


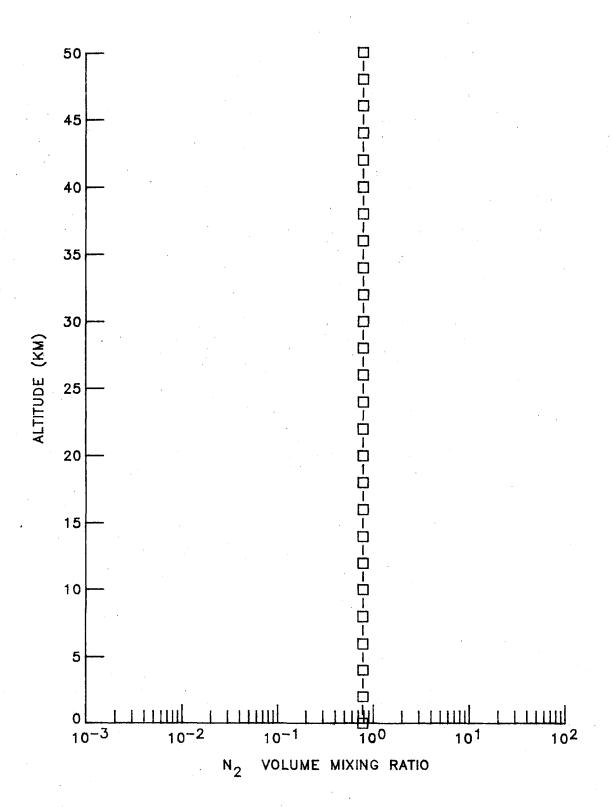


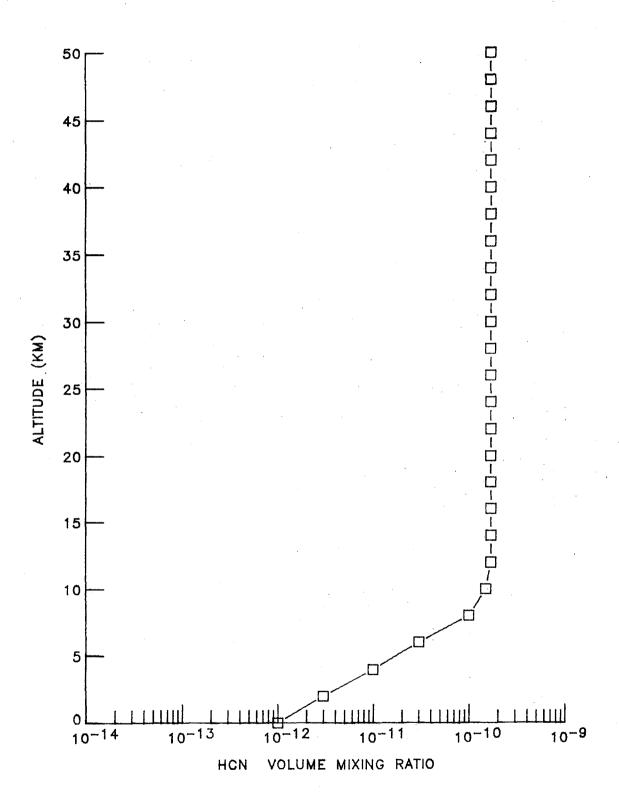


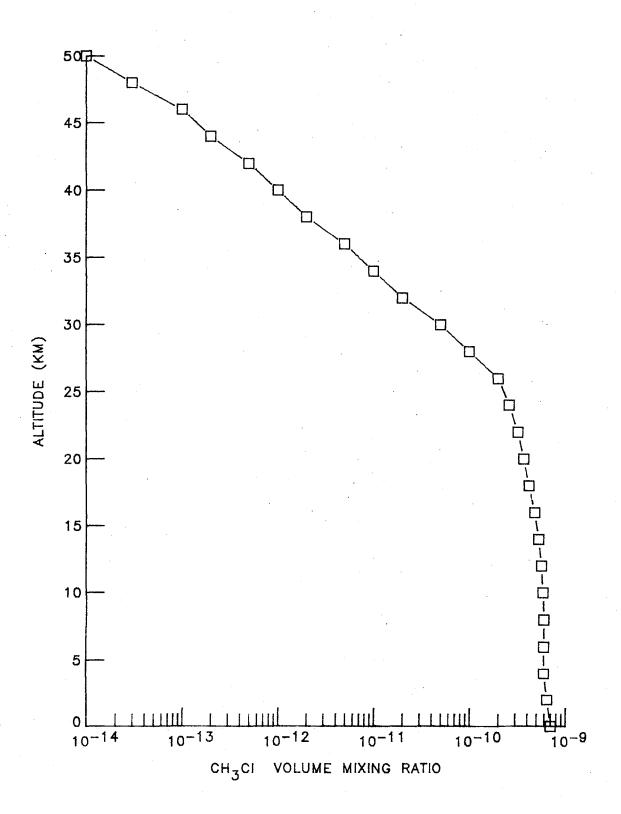


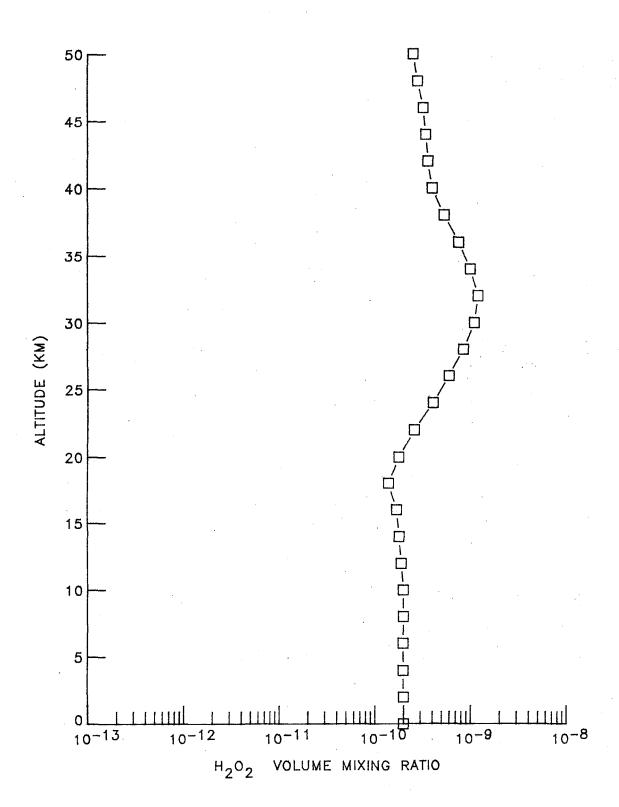


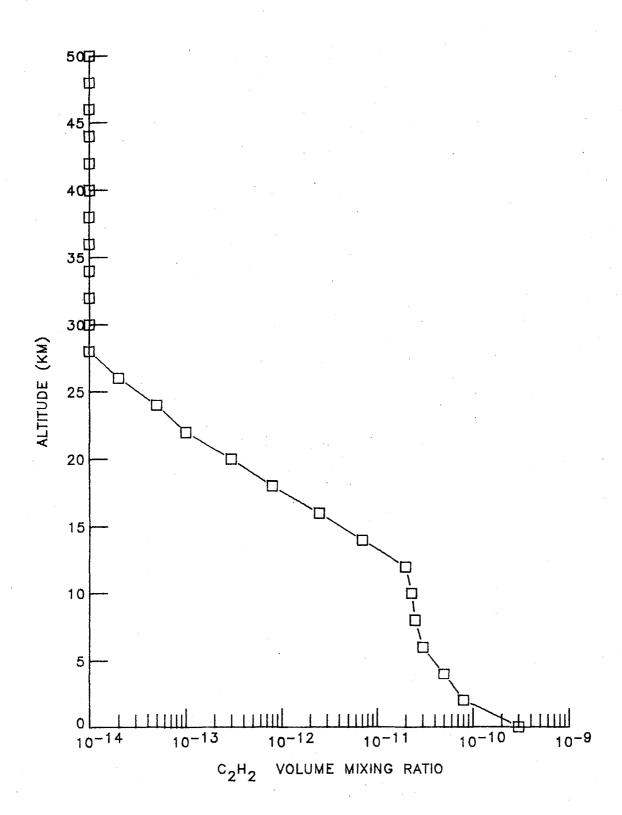


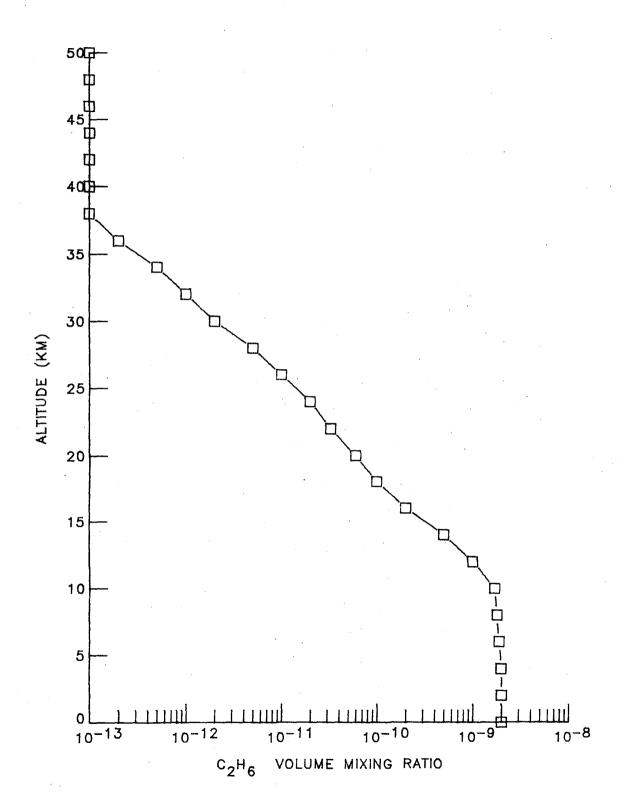


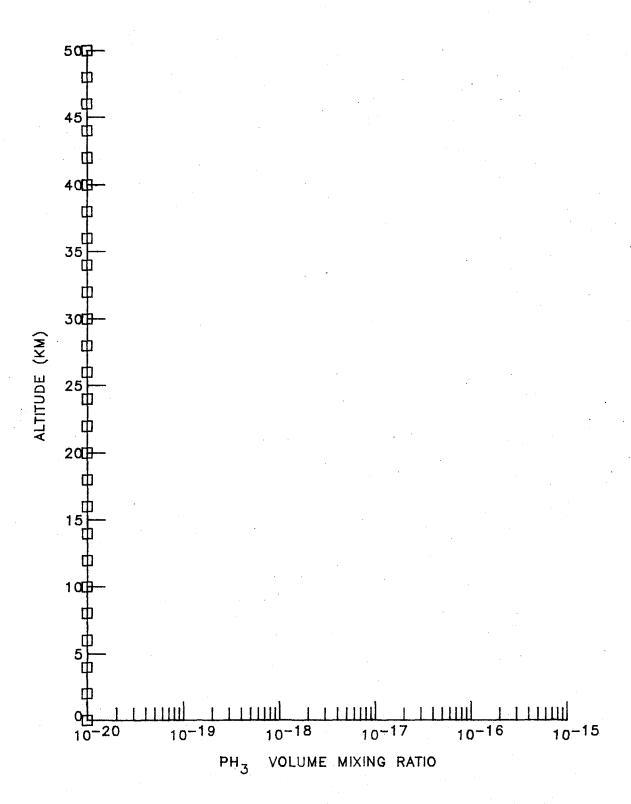


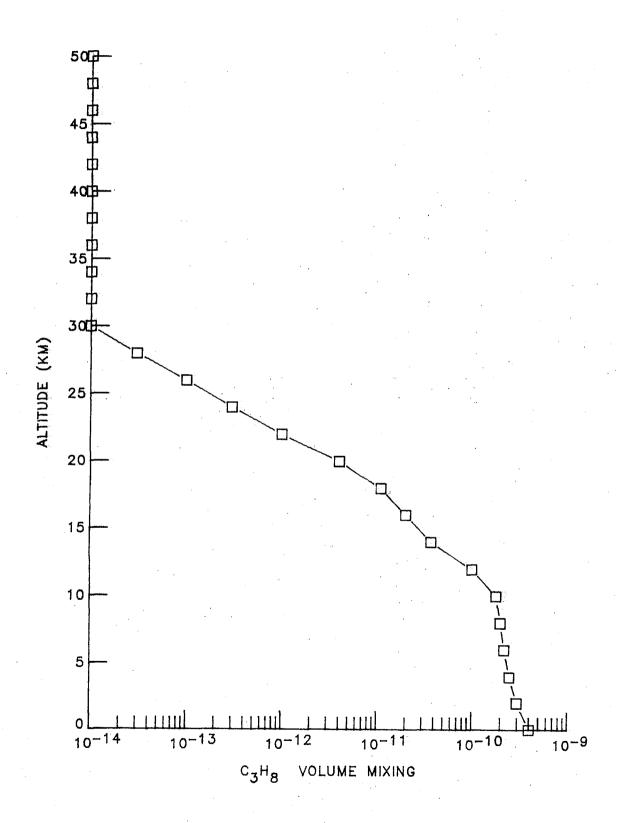


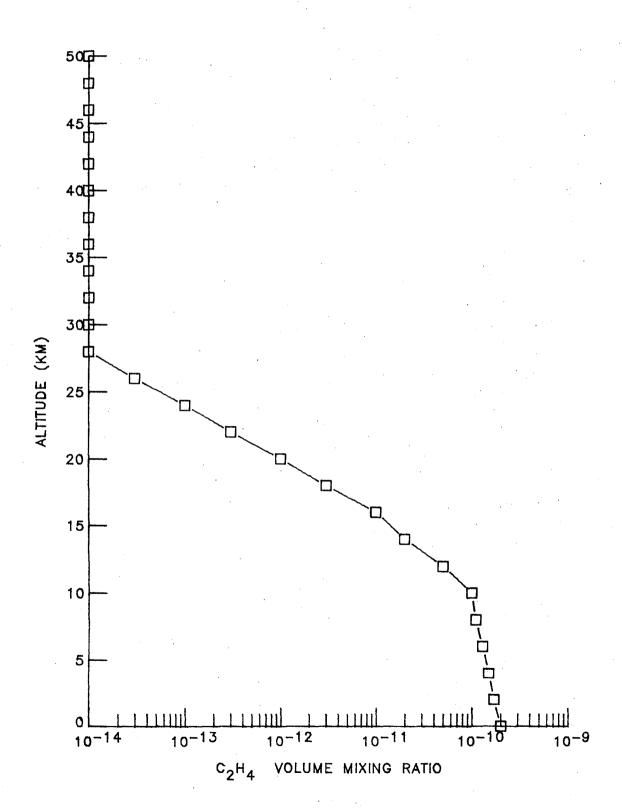


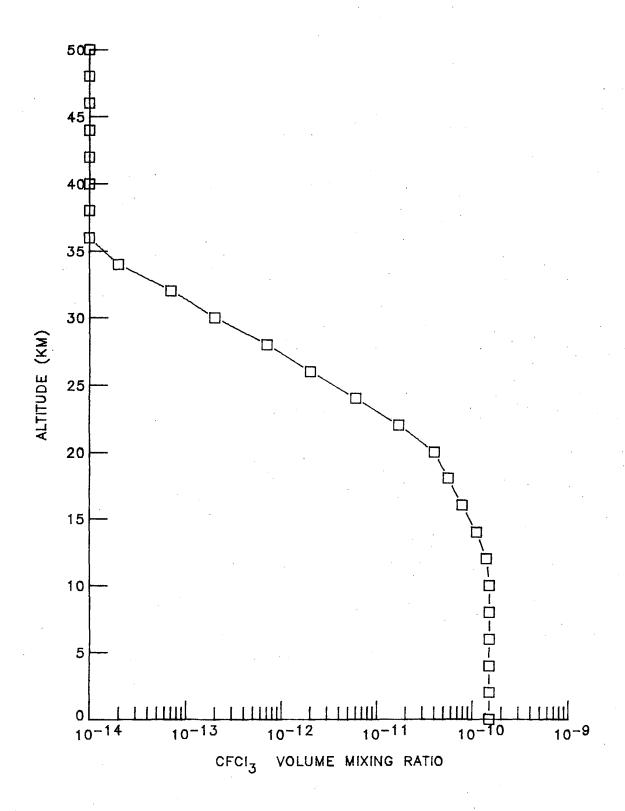


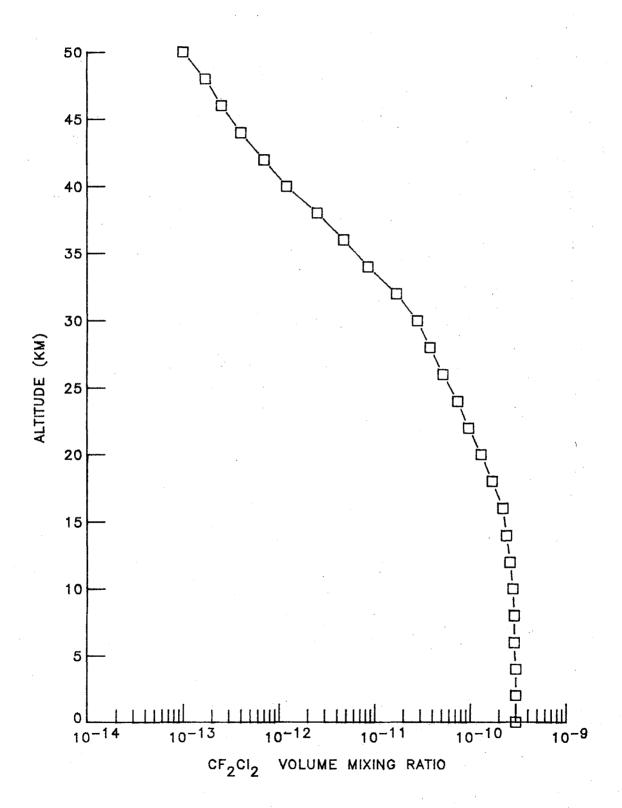


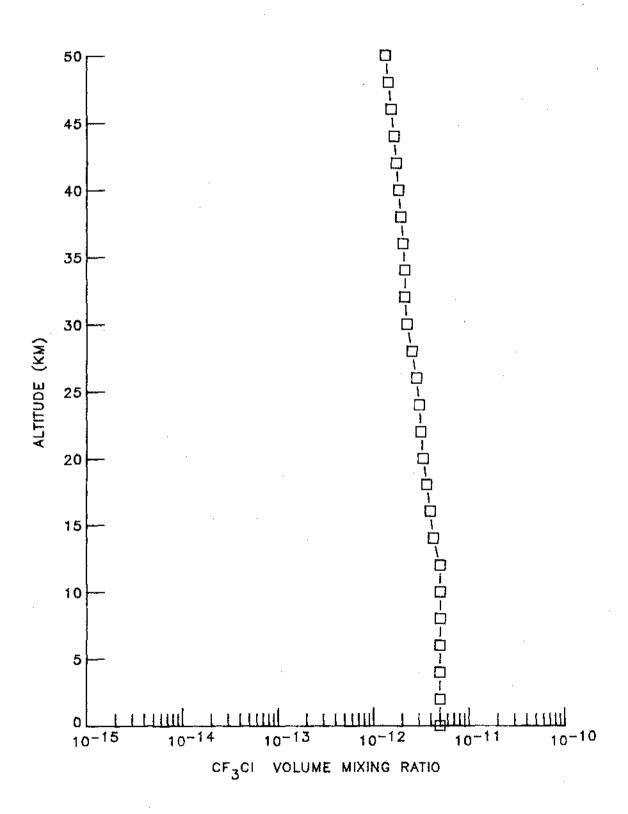


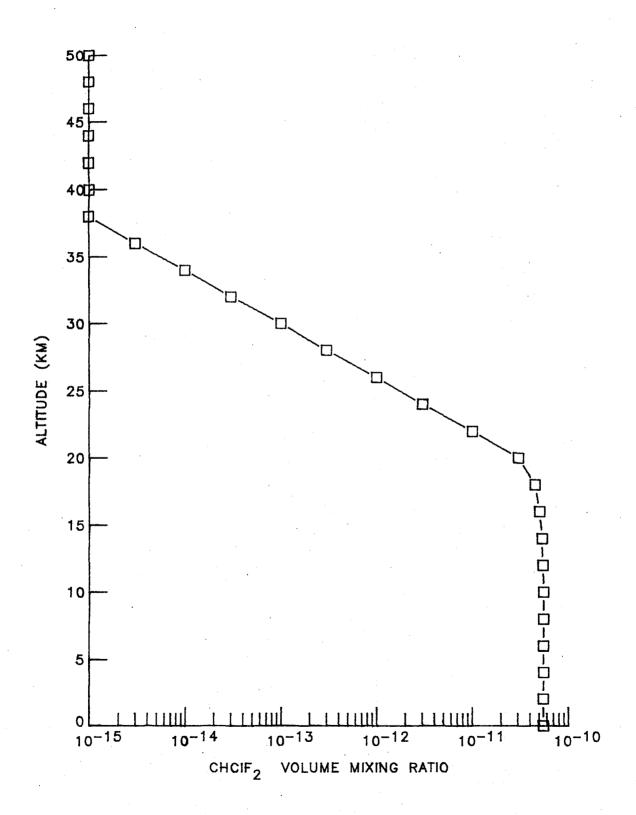


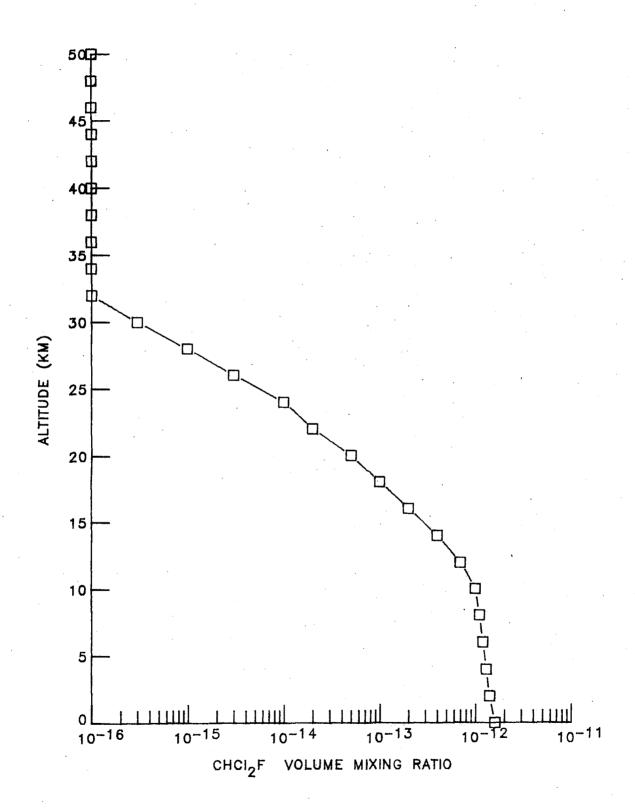


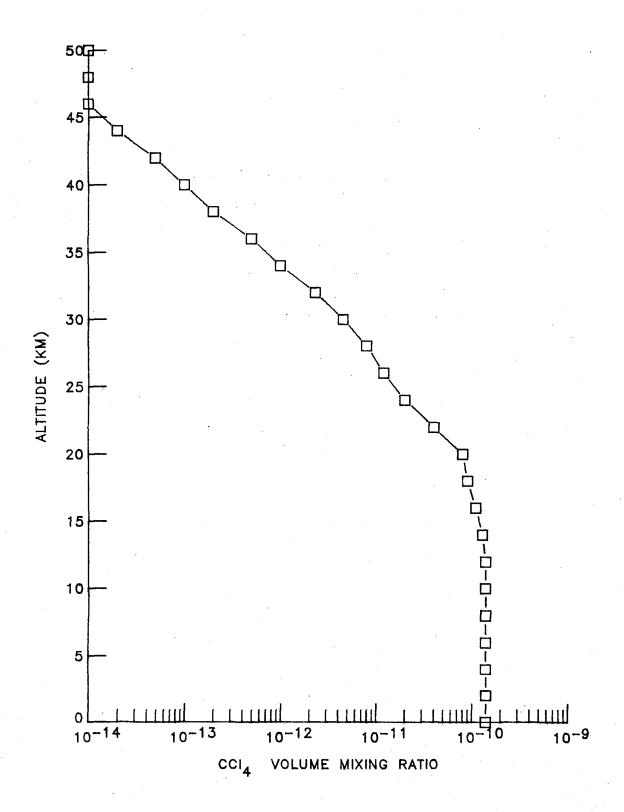


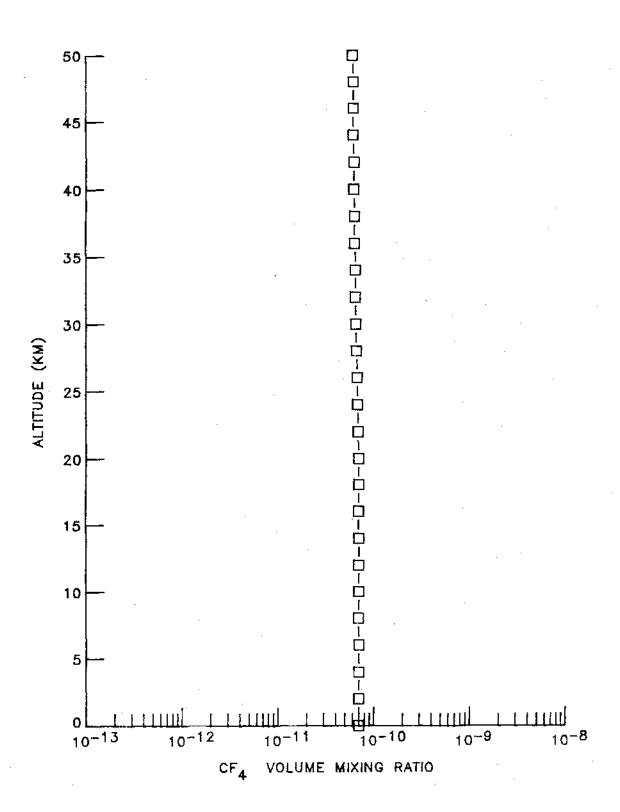


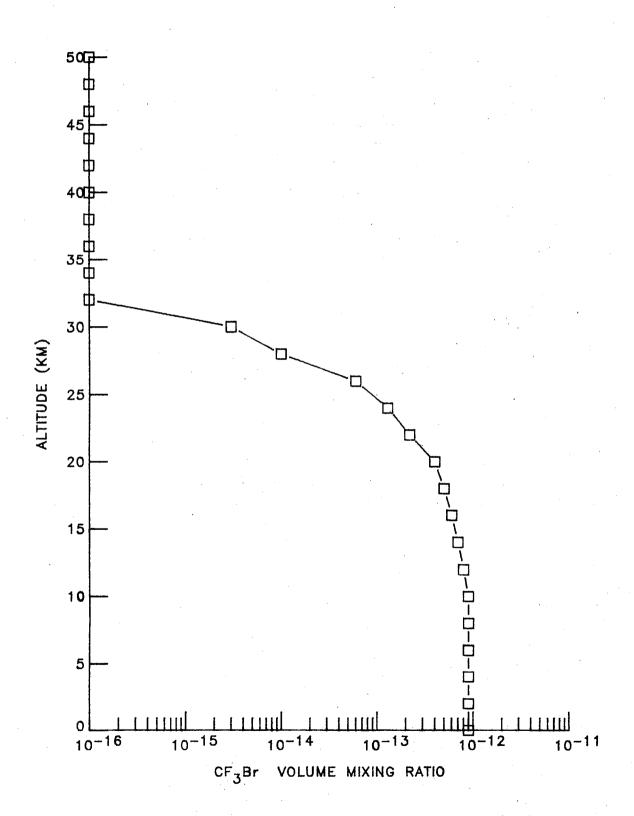


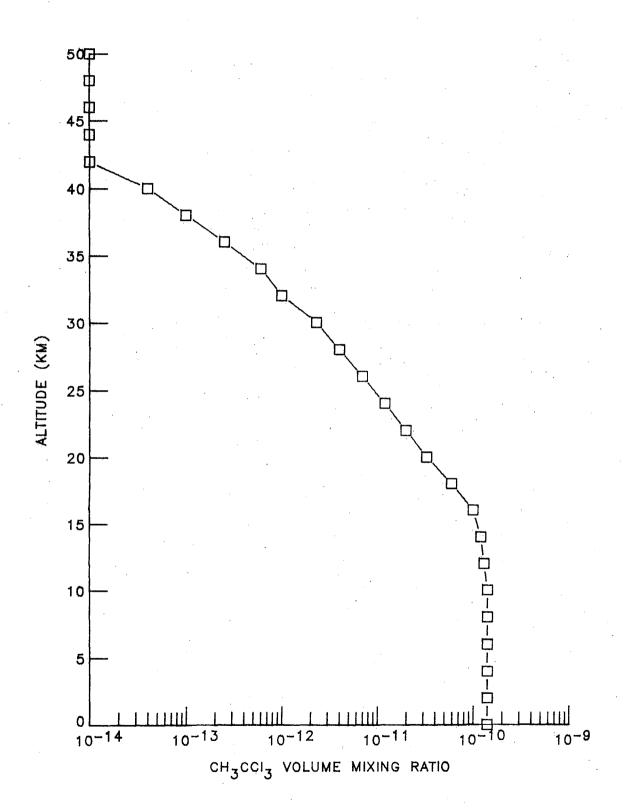


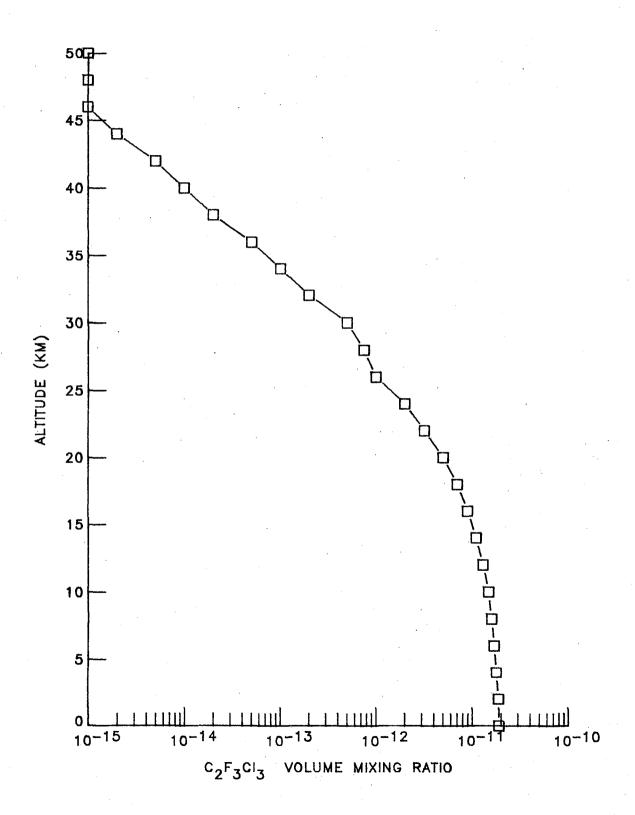


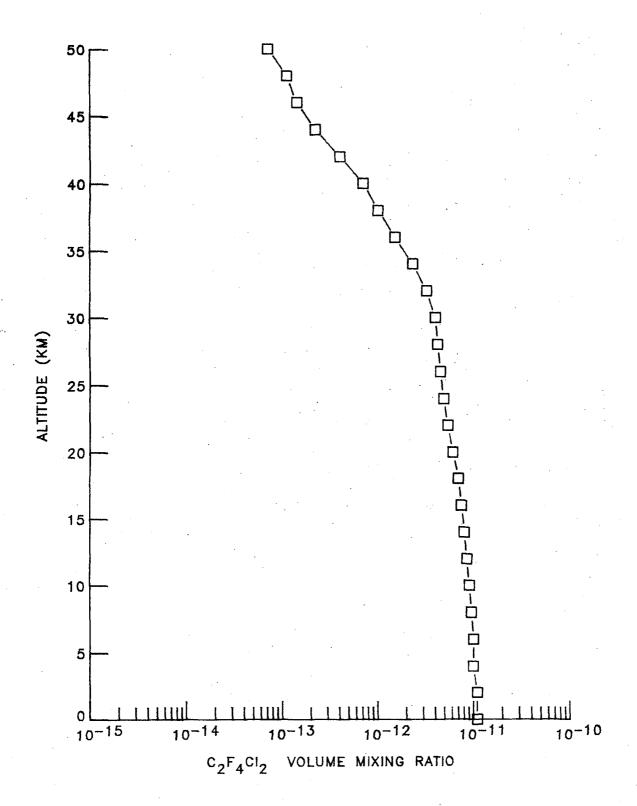


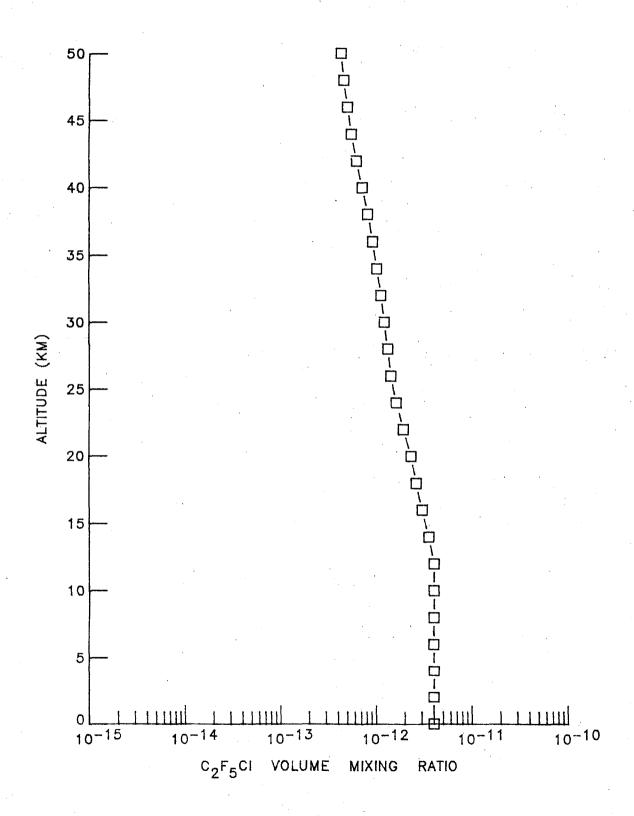


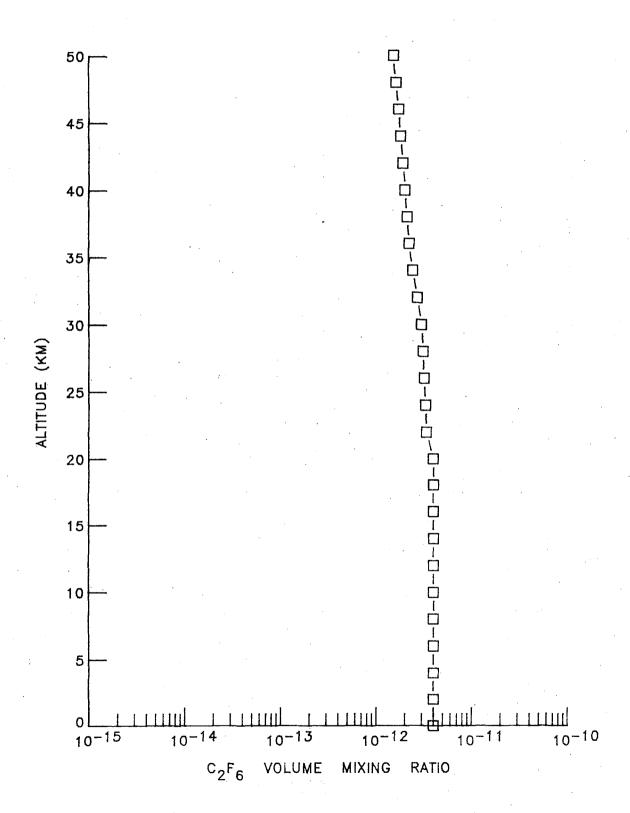


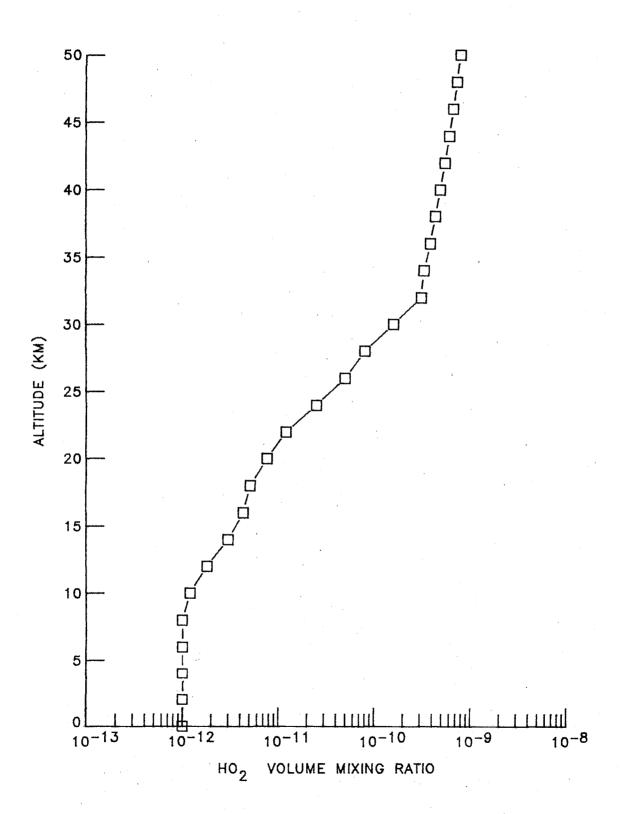


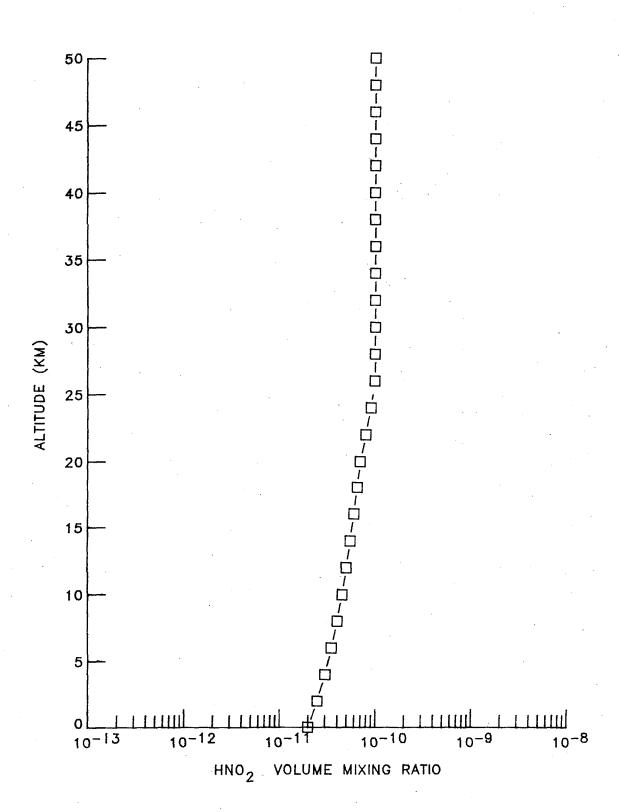


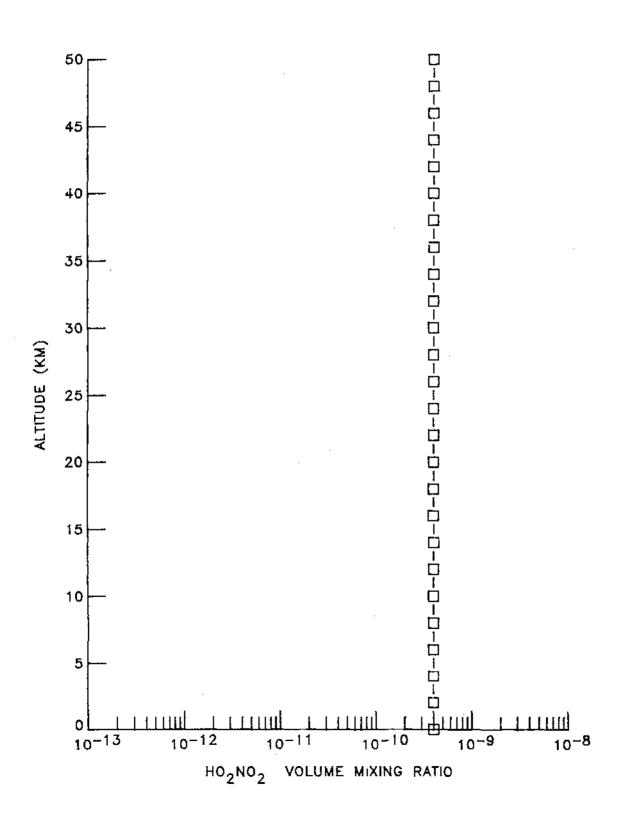


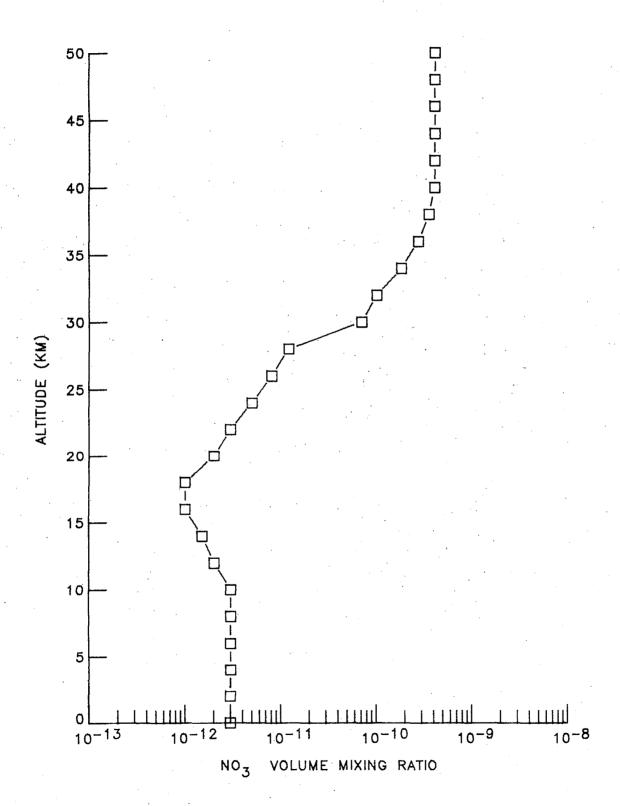


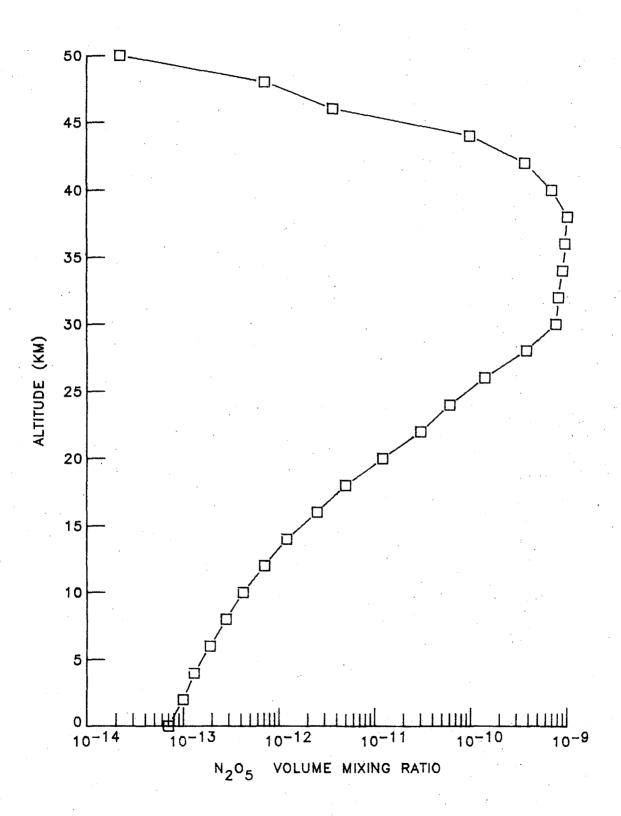


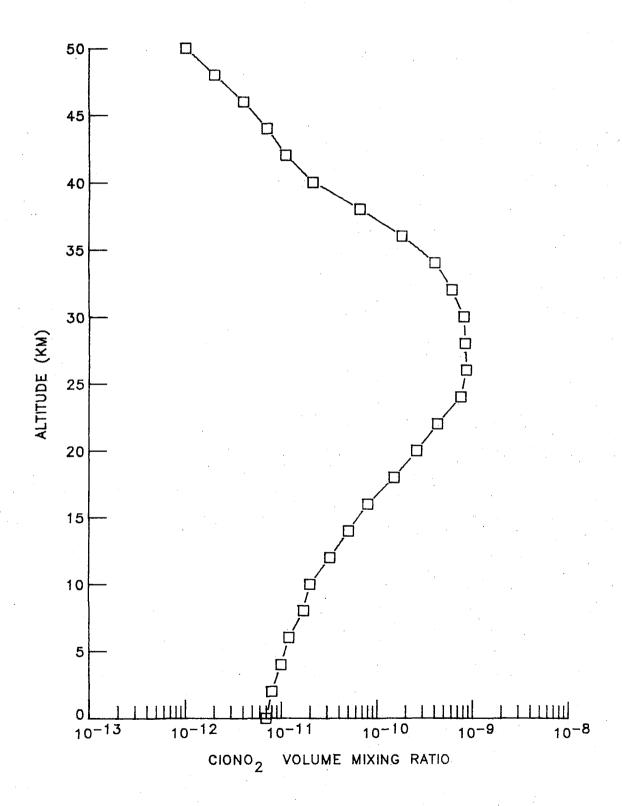


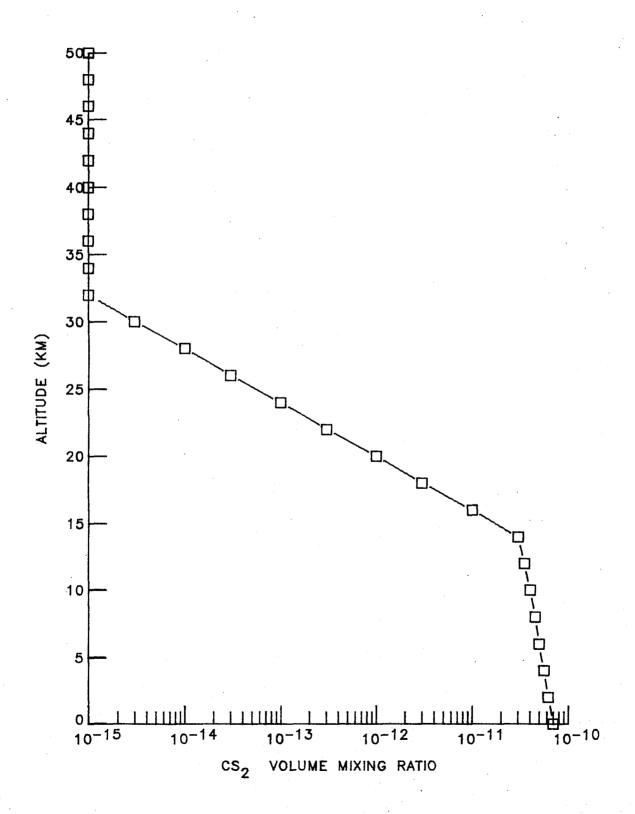


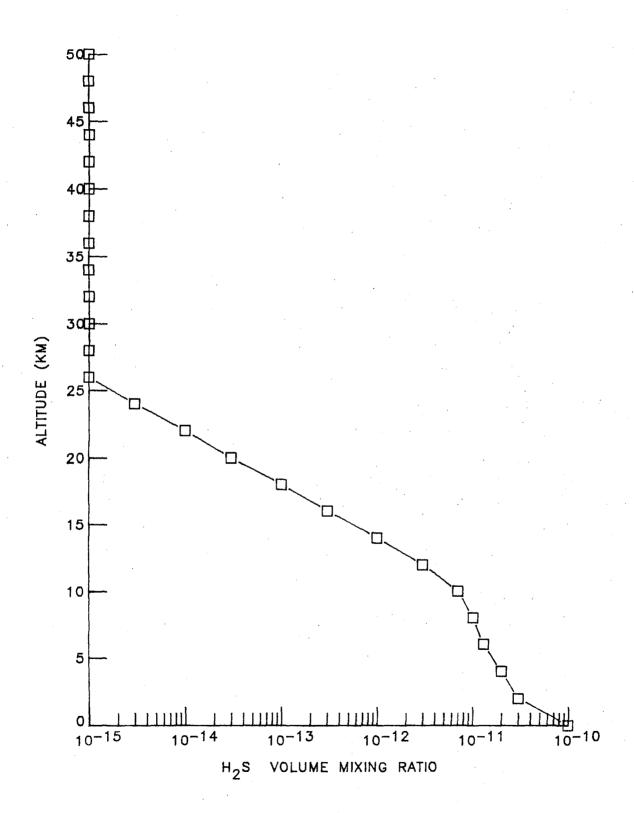


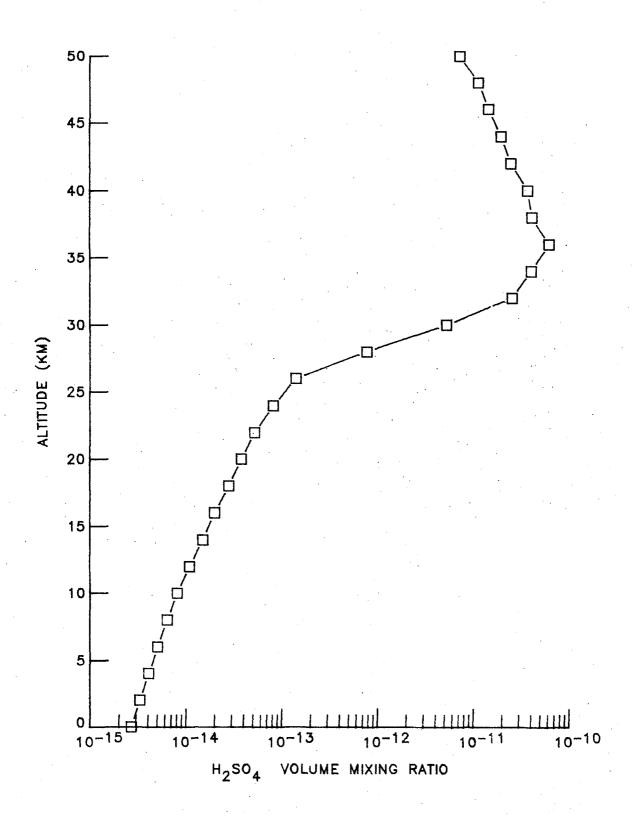












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